Polarized radio emission from extensive air showers measured with LOFAR

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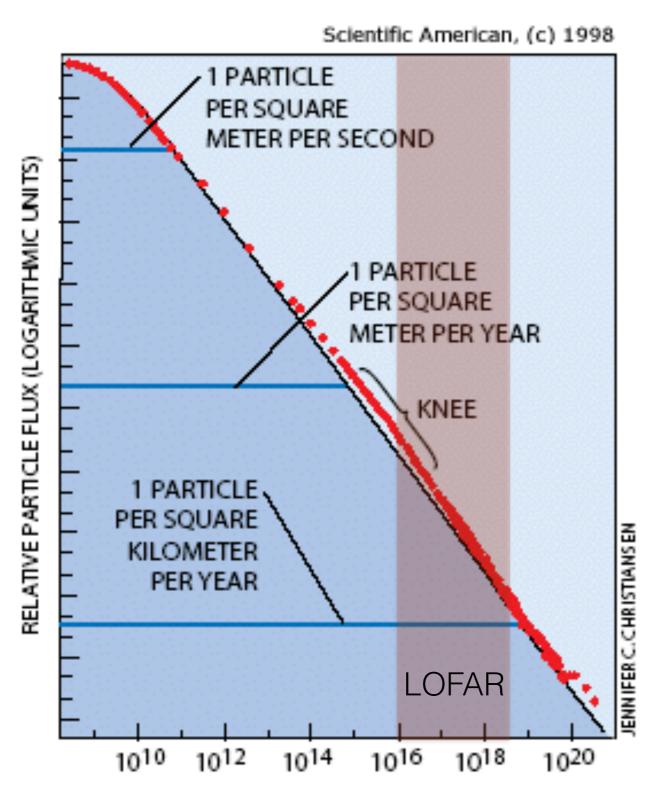
Cosmic Rays Key Science Project

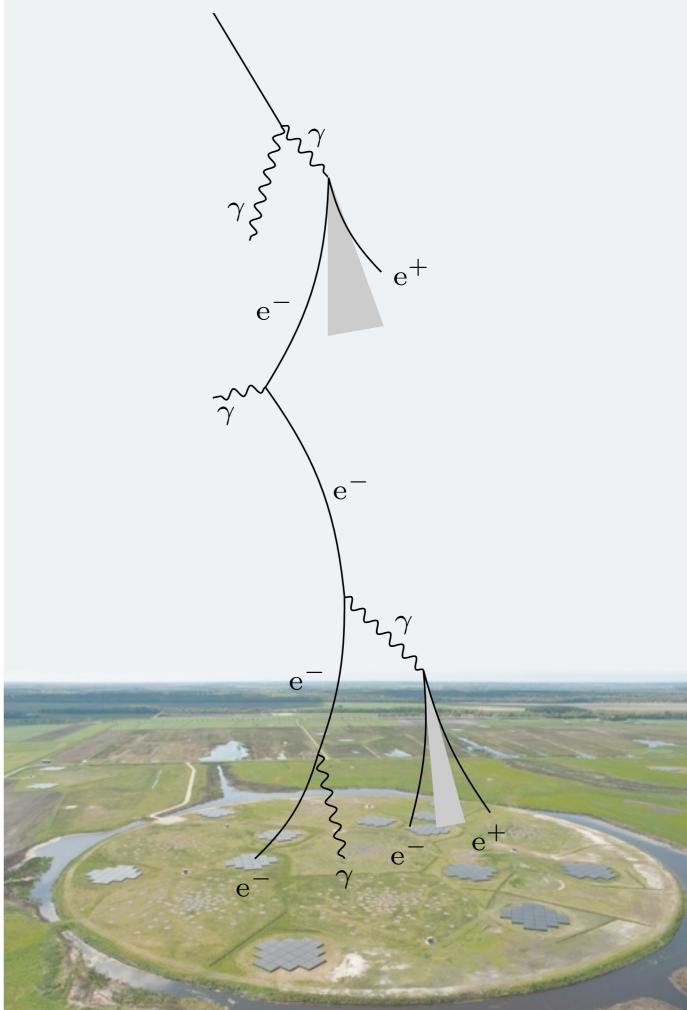
Radboud University Nijmegen, ASTRON, NIKHEF, Max-Planck-Institut für Radioastronomie, Rijksuniversiteit Groningen

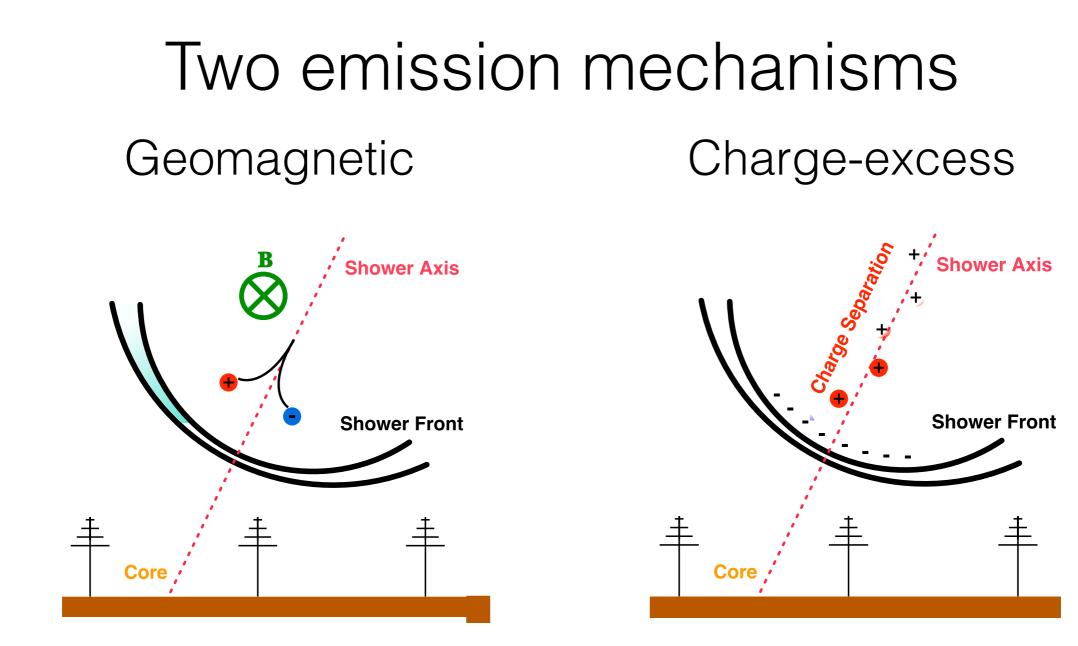


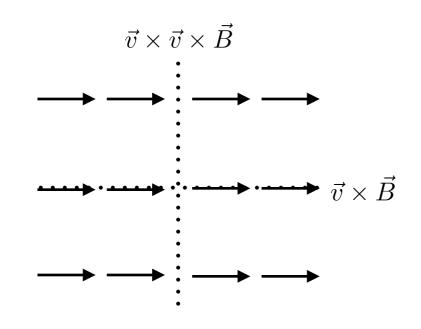
Radboud University Nijmegen

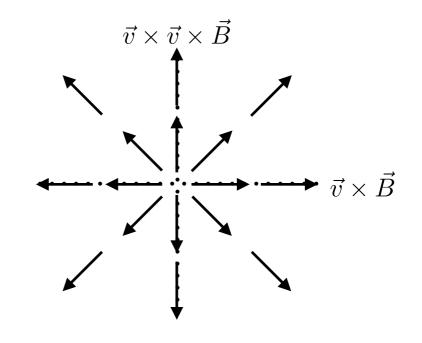
Radio emission from air showers











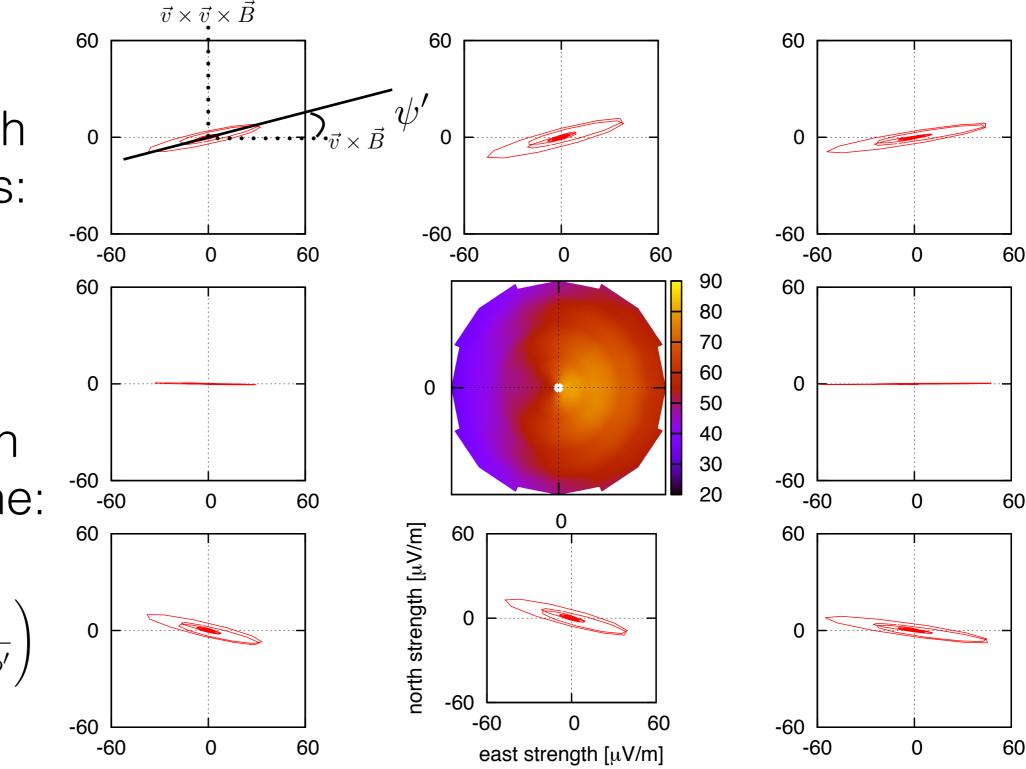
Polarization angle depends on:

relative strength of contributions:

$$a \equiv \sin \alpha \frac{|E_{\rm C}|}{|E_{\rm G}|},$$

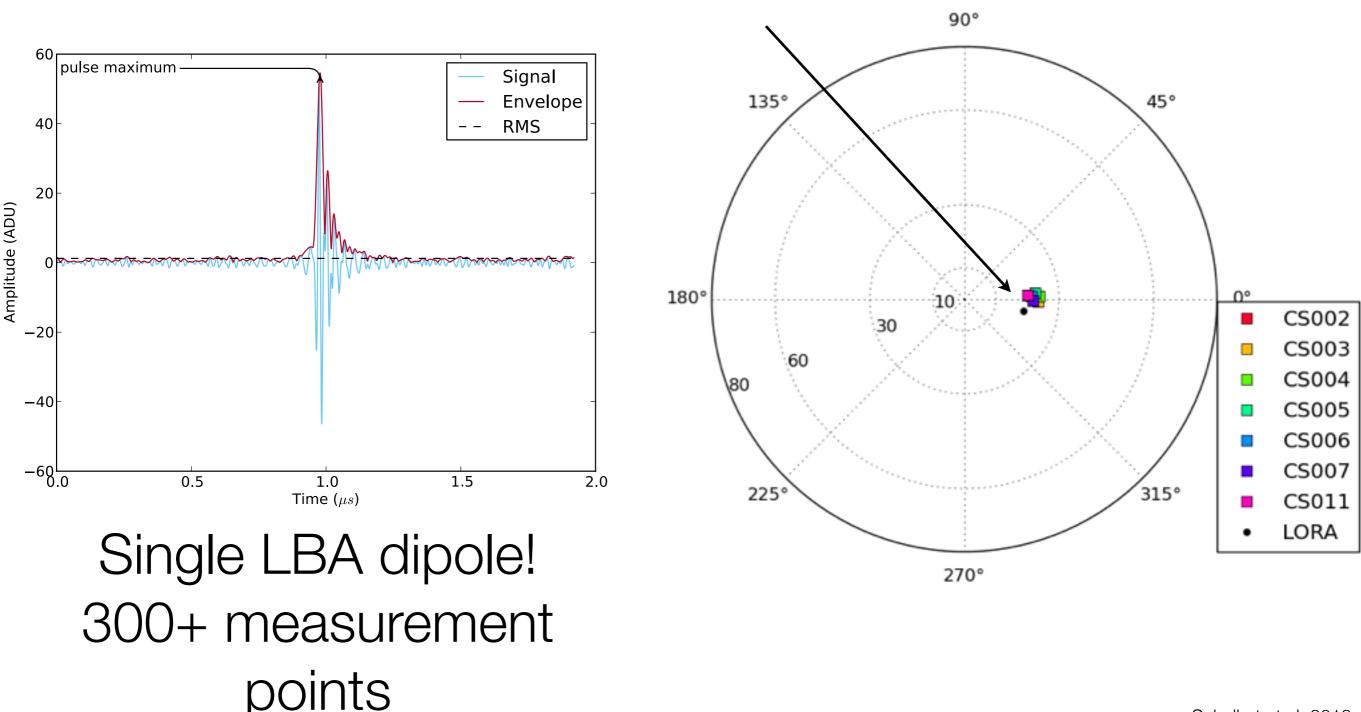
and, position in the shower plane:

$$\psi' = \tan^{-1} \left(\frac{\sin \phi'}{\frac{\sin \alpha}{a} + \cos \phi'} \right)$$



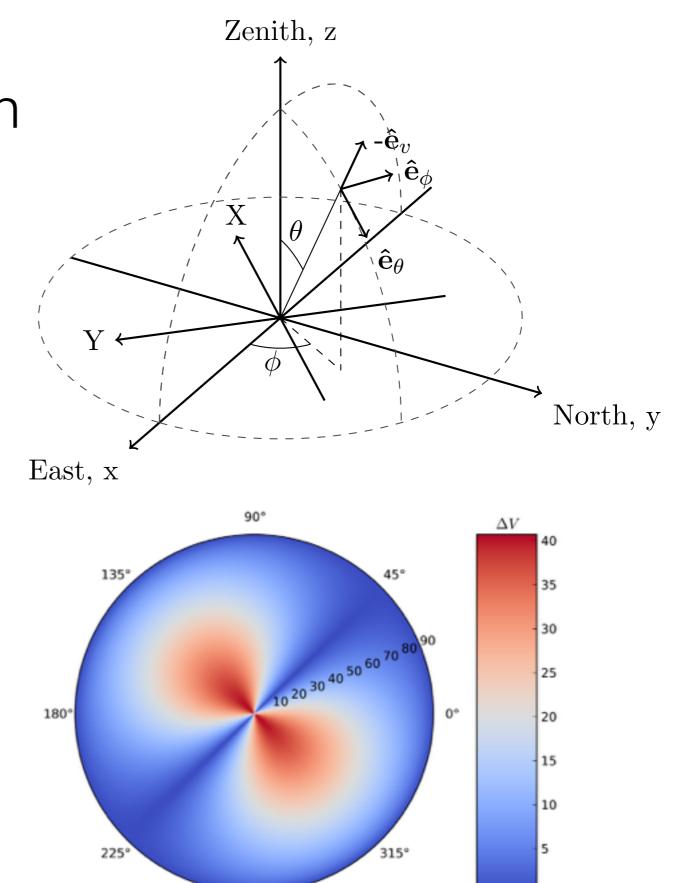
Automated detection pipeline (400+ showers measured) Typical event:

Particle Detector & Radio Direction Reconstruction



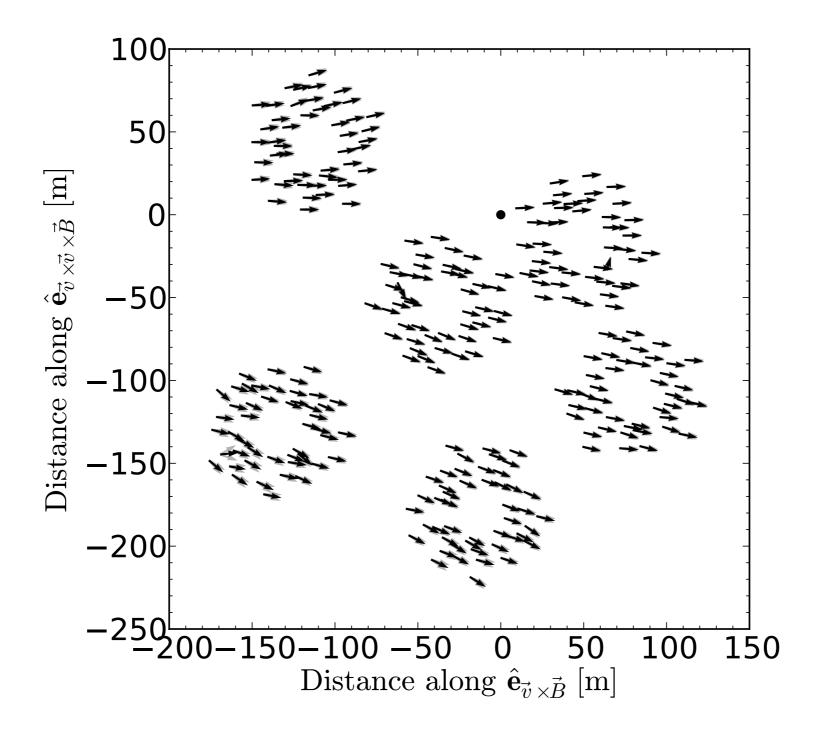
Use antenna simulation to reconstruct polarization

- WIPLD EM-simulation + electronics model (by Michel Arts @ ASTRON)
- Complex direction and frequency dependent gain per polarization per dipole
- Interpolate directly to get 2x2 complex Jones matrix for pulse direction
- Invert and multiply to get E(t)

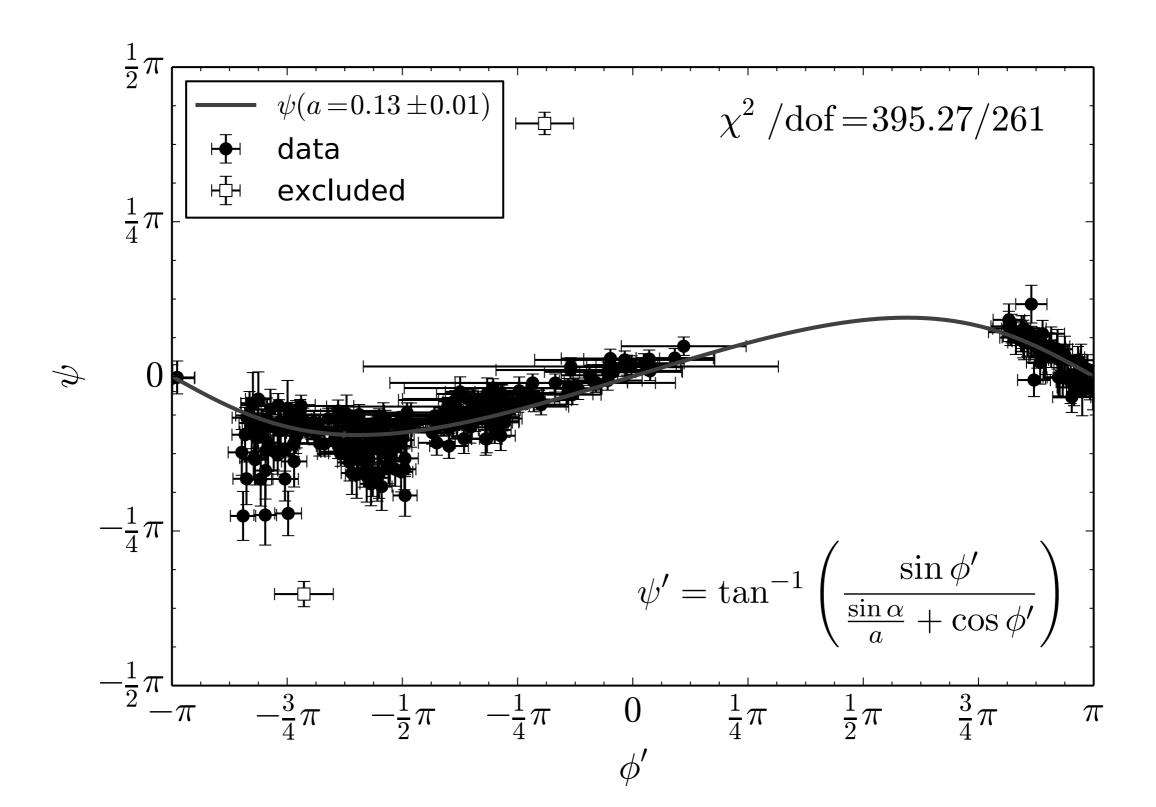


270°

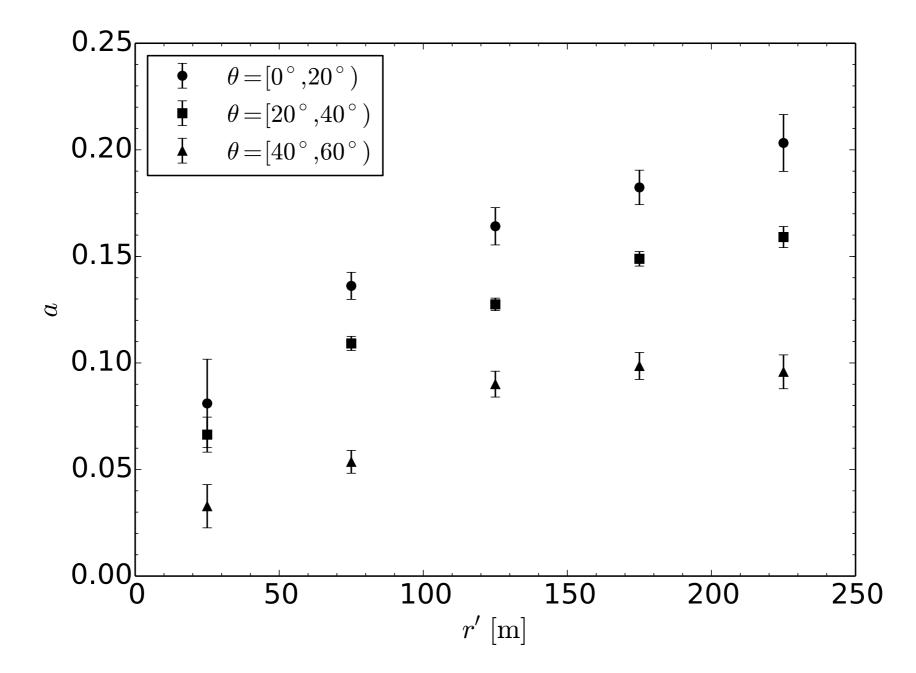
Observed polarization pattern



Charge-excess fraction

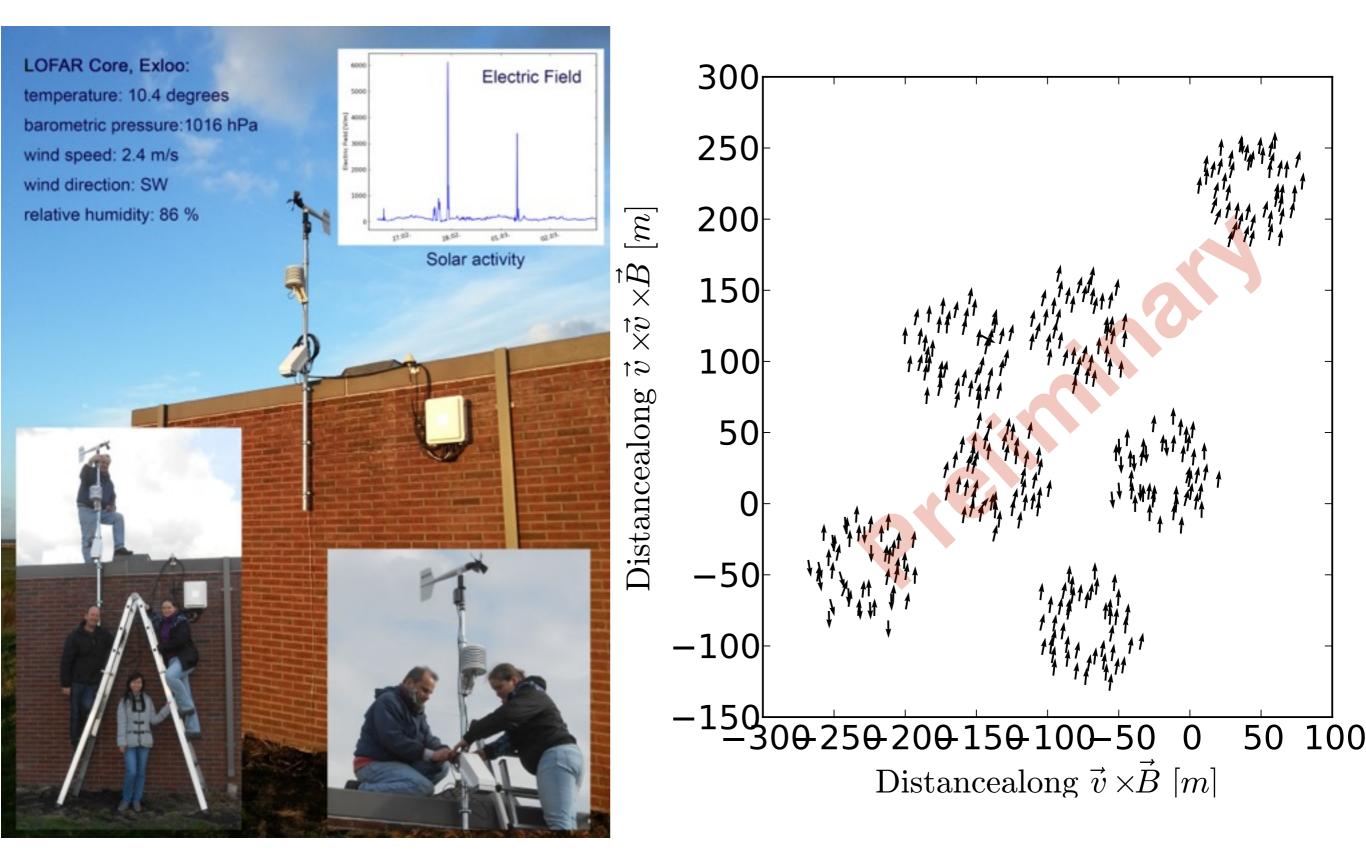


Confirmed predicted dependence on zenith angle and radial distance



(Schellart et al. 2014, submitted to JCAP) Made possible by high antenna density at LOFAR

Atmospheric electric fields



Conclusions

- Polarization measurements can be used to reconstruct relative contributions of emission mechanisms to air shower radio emission
- Uniquely high antenna density offered by LOFAR allows for first time confirmation of predicted radial and zenith angle dependence of charge-excess fraction (Schellart et al. 2014, submitted to JCAP)
- Strong atmospheric electric fields (thunderstorms) alter the observed polarization pattern